

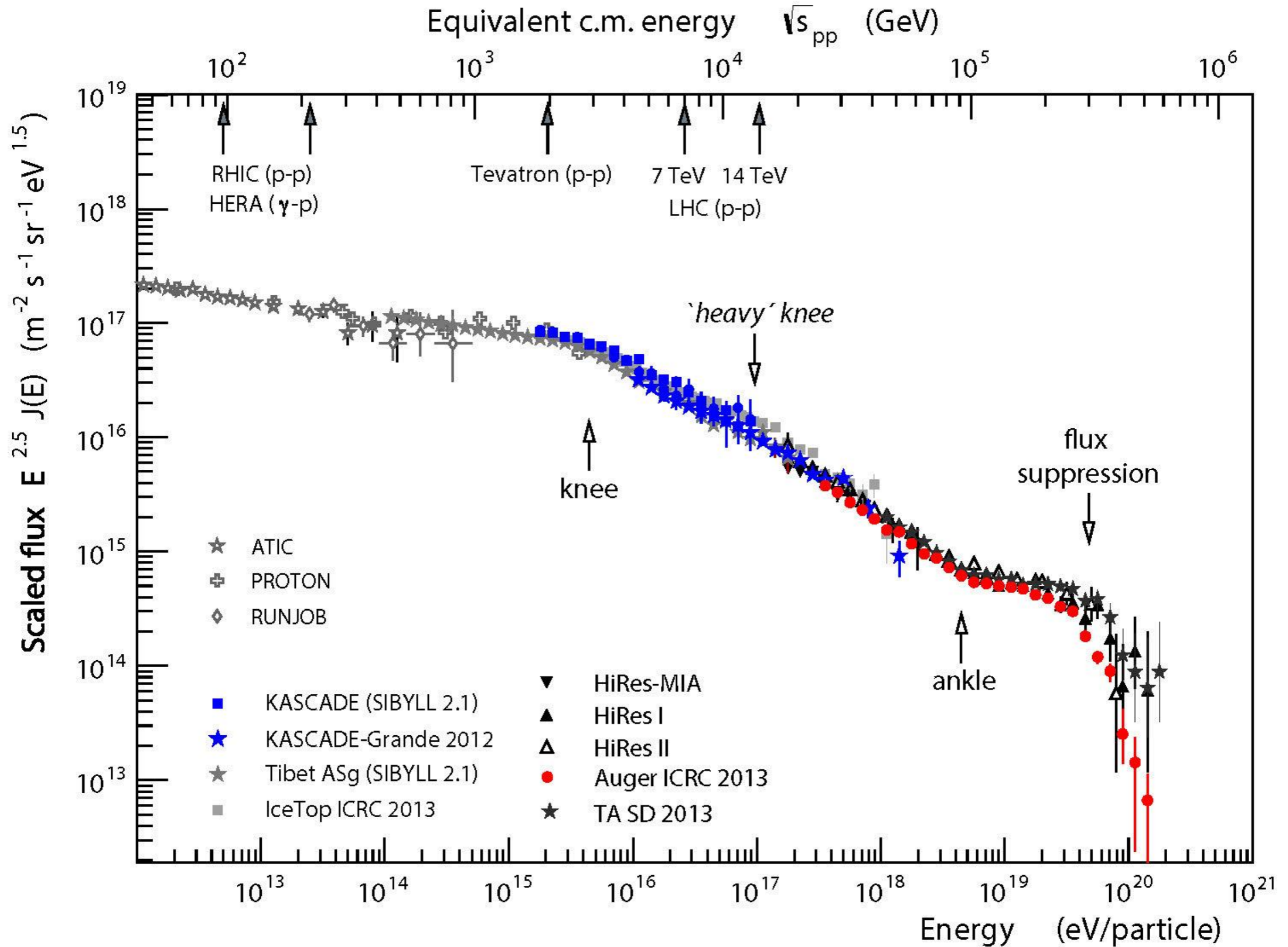
APPEC – Town Meeting
Roadmap (“Considerations”) Discussion
(High-Energy) Cosmic Rays
Paris, 6-7 April 2016



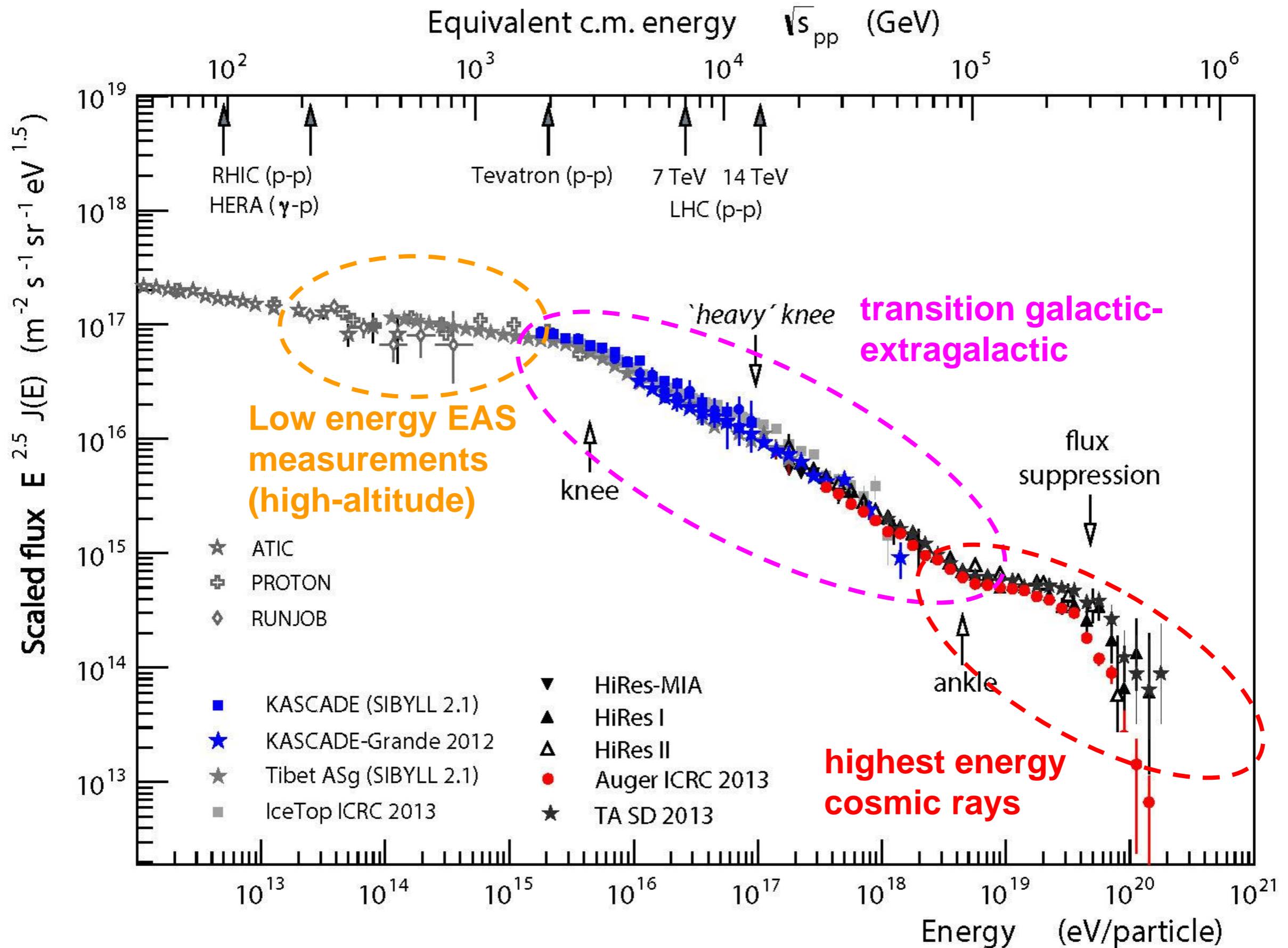
Andreas Haungs
Karlsruhe Institute of Technology
haungs@kit.edu

Petr Tinyakov
Université Libre de Bruxelles
petr.tiniakov@ulb.ac.be

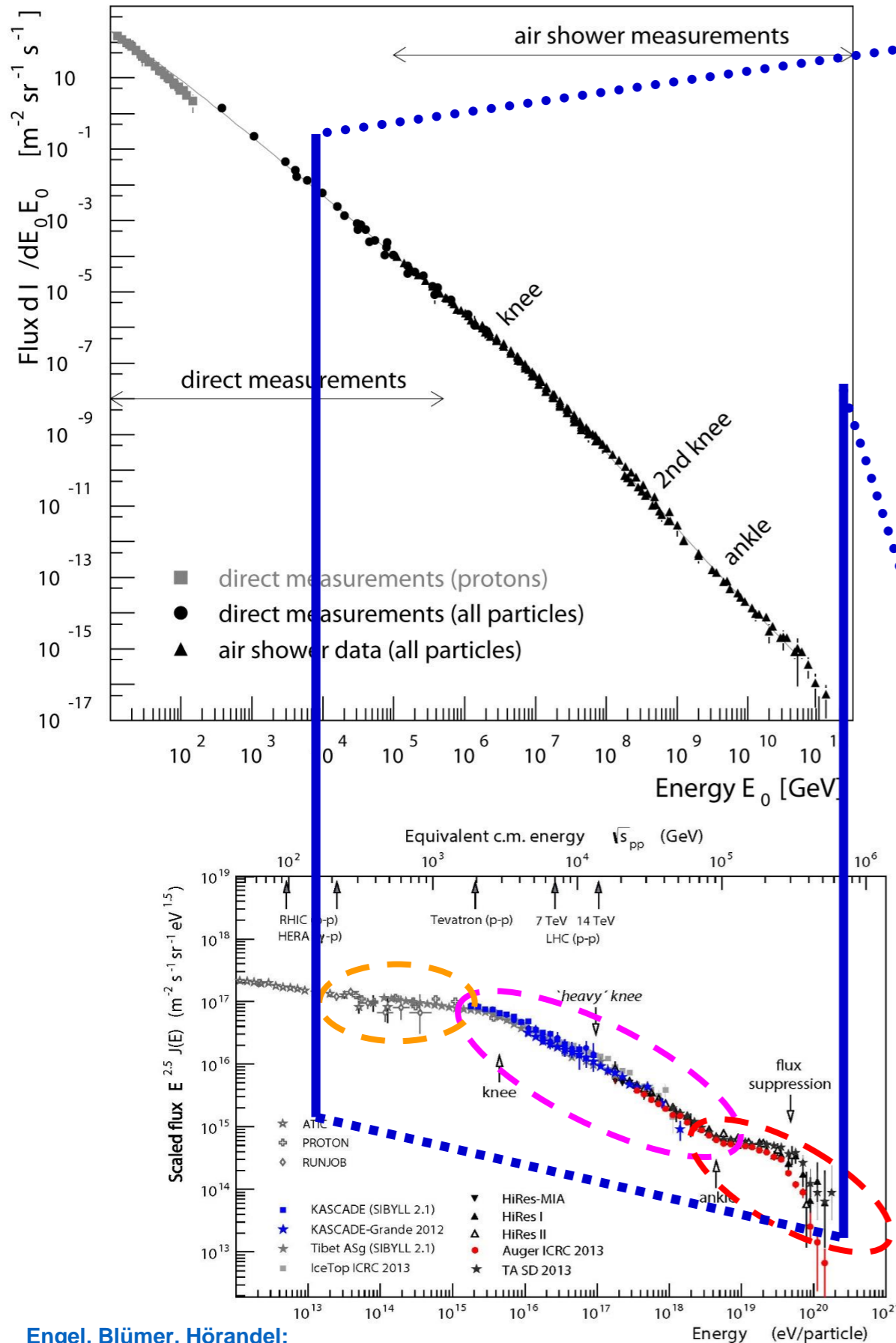
Energy spectrum



Energy spectrum



Questions / Challenges of the experiments



Spectral indices?

Antimatter?

Composition / Isotopes?

Overlap direct-indirect measurements?

Hadronic interaction models?

Fine-structures in spectrum?

End of Galactic Spectrum?

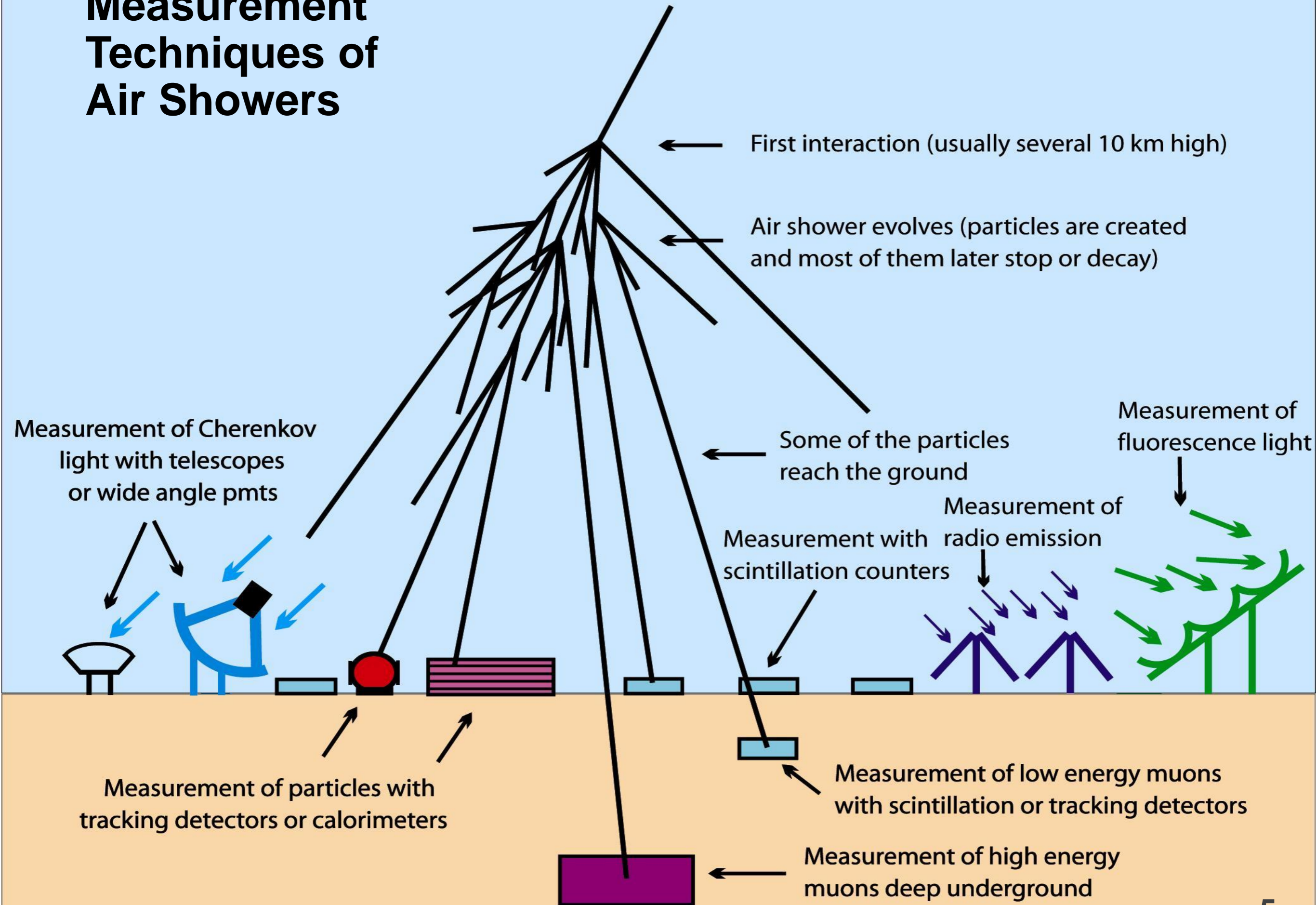
Composition?

GZK / maximum acceleration?

Anisotropy?

...and many more!

Measurement Techniques of Air Showers



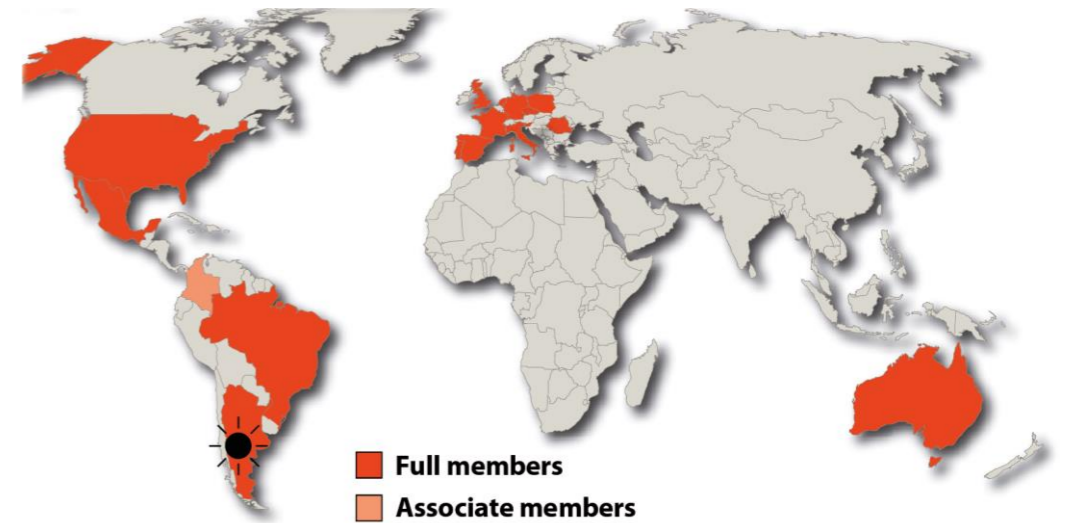
Pierre Auger Observatory



Pierre Auger Observatory: Results

one decade ago:

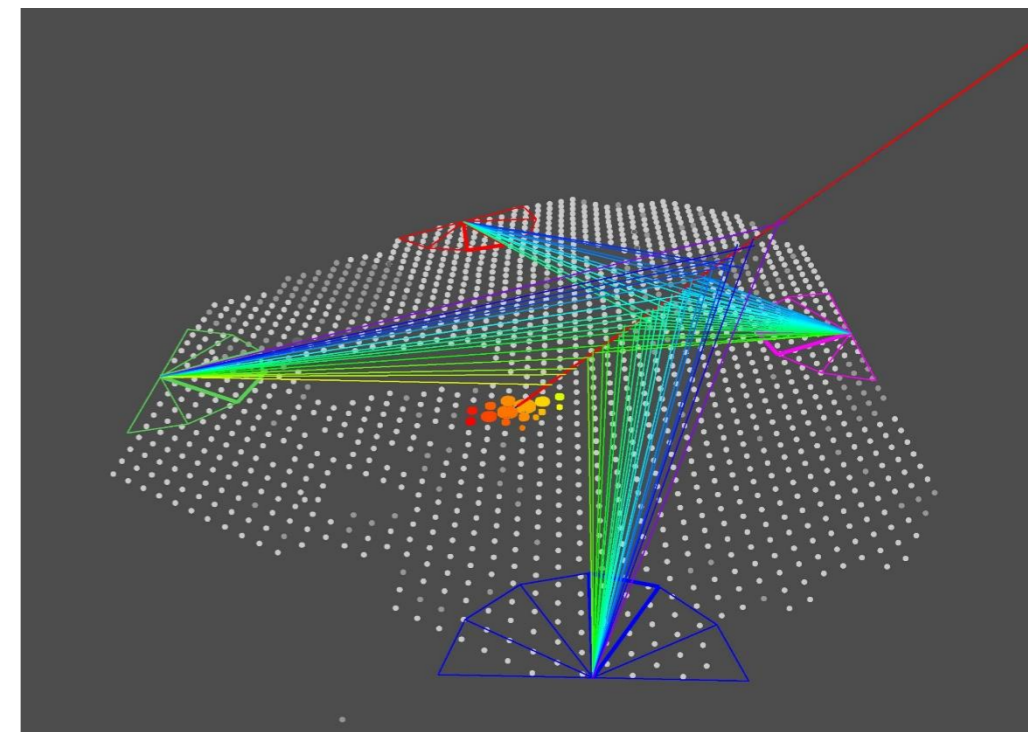
- not known whether flux suppression exists
- composition expected to be protons
- no anisotropies observed



~450 collaborators; 92 institutions, 17 countries

today:

- flux suppression (50 EeV) beyond any doubt
- anisotropies (large scale) seen
- evidence to become heavier
- cosmogenic γ and ν are constrained
- particle physics performed $\sigma(pp)$ @57TeV cms
- atmospheric effects, LIV studies,
-

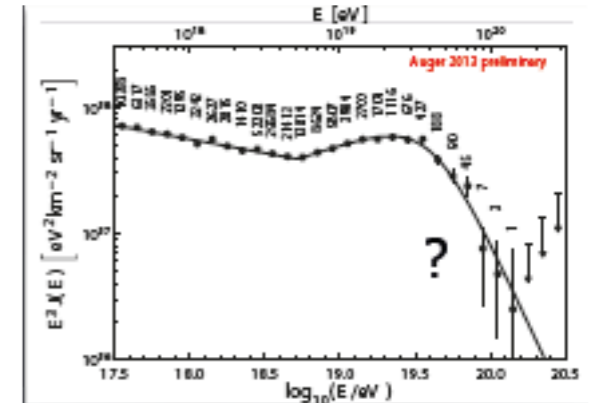


- but, flux suppression mechanism unclear
- but, no sources of UHECR identified - yet

Auger-upgrade (AugerPrime) Science Goals

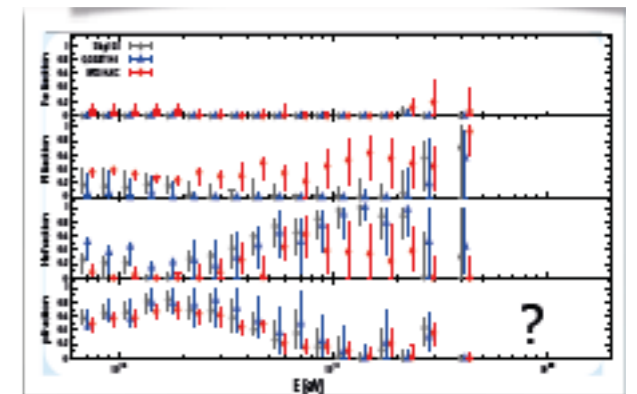
1. Elucidate the origin of the flux suppression, i.e. GZK vs. maximum energy scenario

- fundamental constraints on UHECR sources
- galactic vs extragalactic origin
- reliable predictions of GZK ν - and γ -fluxes



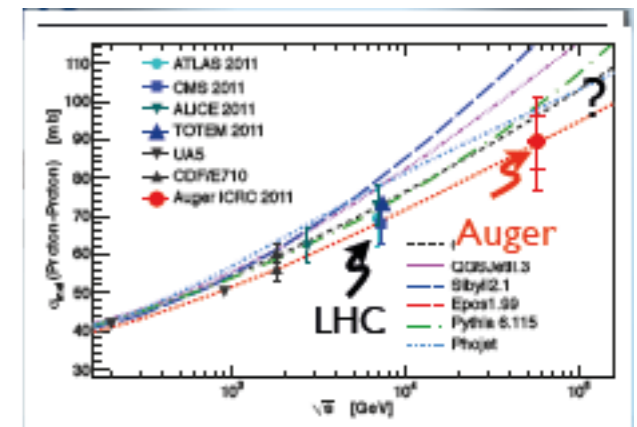
2. Search for a flux contribution of protons up to the highest energies at a level of ~ 10%

- proton astronomy up to highest energies
- prospects of future UHECR experiments



3. Study of extensive air showers and hadronic multi-particle production above $\sqrt{s}=70$ TeV

- particle physics beyond man-made accelerators
- derivation of constraints on new physics phenomena



Pierre Auger Observatory is in place to address all these questions now

Auger upgrade (AugerPrime)

The Pierre Auger Observatory Upgrade

Preliminary Design Report



April 10, 2015

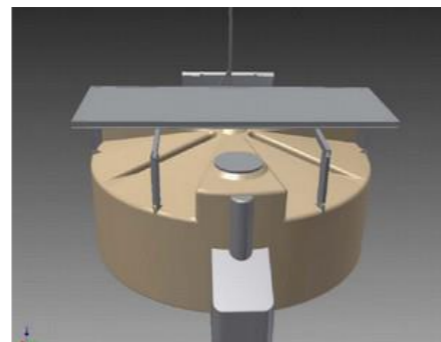
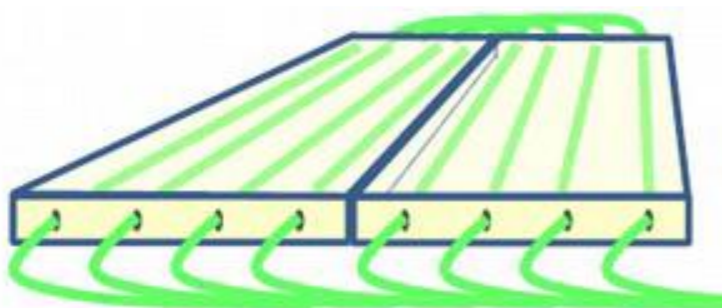
Organization: Pierre Auger Collaboration
Observatorio Pierre Auger,
Av. San Martín Norte 304,
5613 Malargüe, Argentina



194 pages

Status:

- positively evaluated by International Advisory Committee
- endorsed by International Finance Board
- R&D well advanced, prototypes running
- 11/2015: International Agreement
- signed for operation into 2025
- engineering array 08/2016
- construction 01/2017 - 2018
- data taking into 2025
- costs: 12.5 M€
- funding: ~50% already committed



Telescope Array



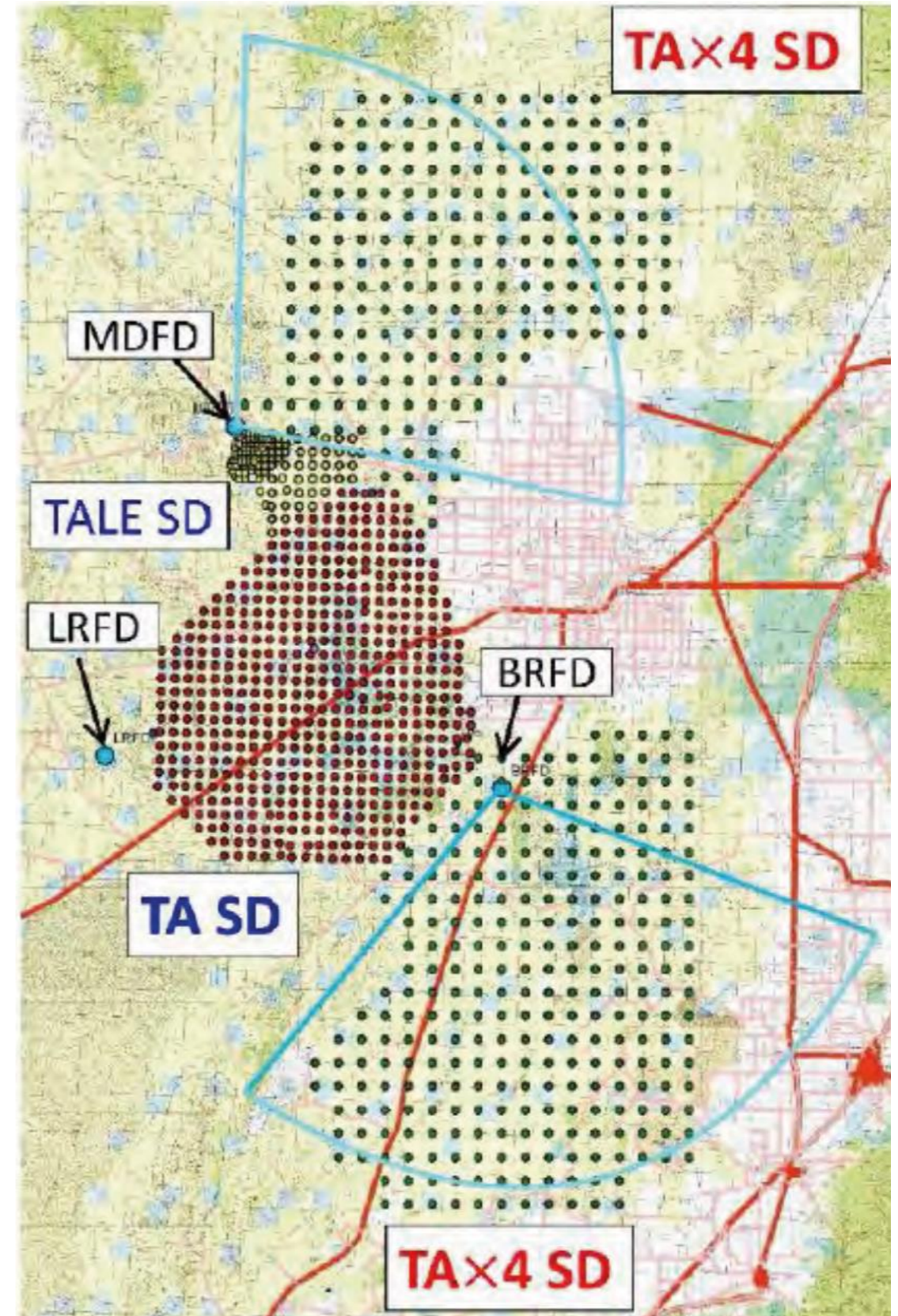
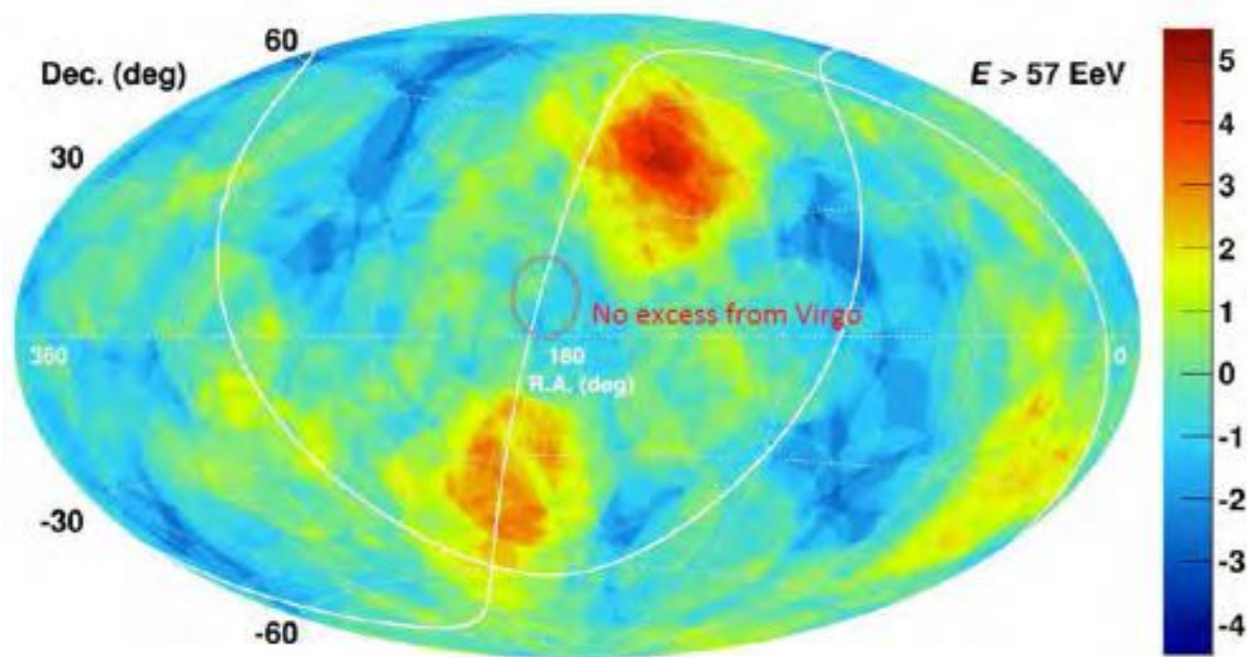
Telescope Array

TAx4:

- SD: 700 → 2800 km²
- Hybrid: x3 acceptance
- Optimized for above ~50 EeV

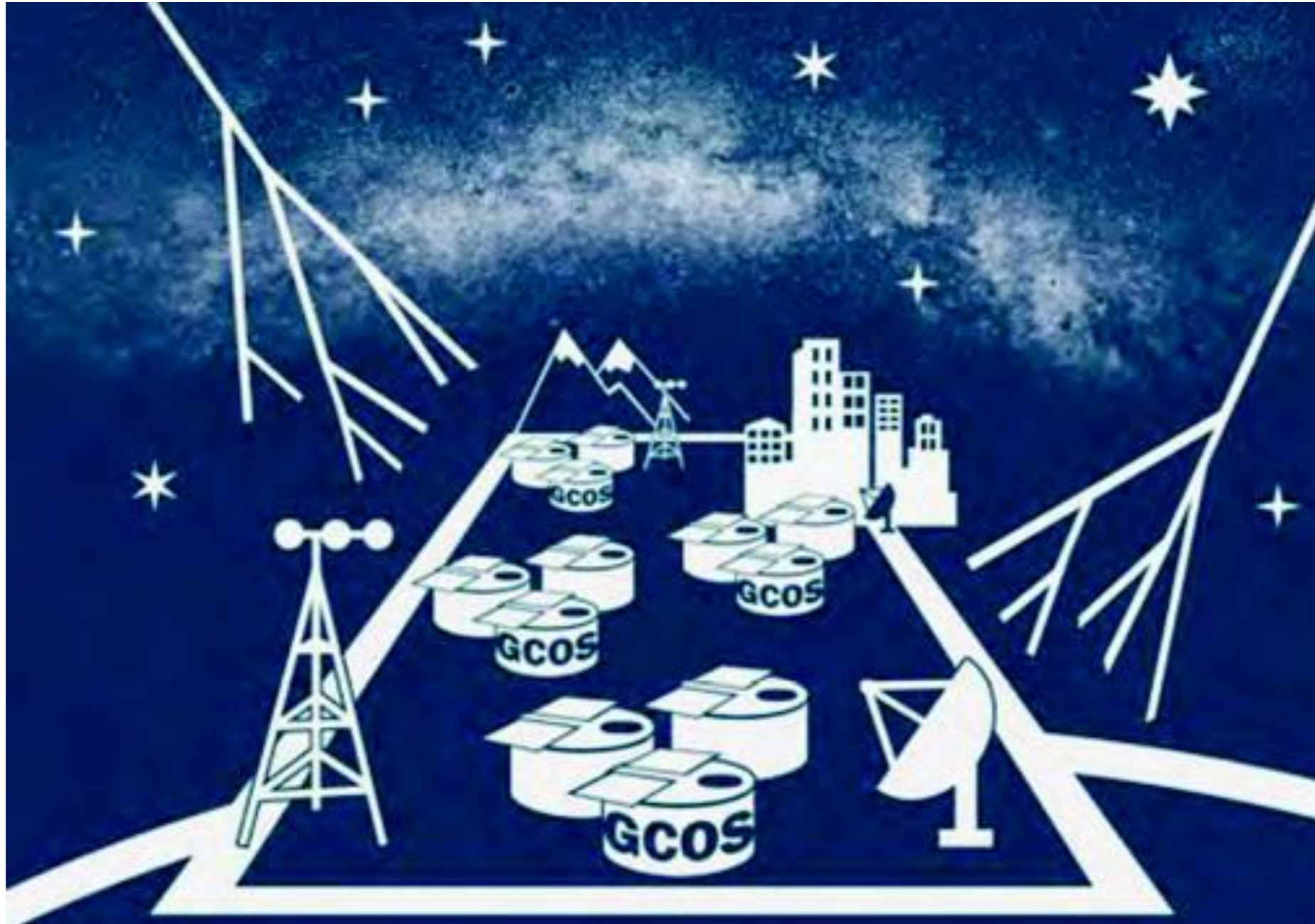
Goal:

Get 20 TA-years by 2019
to look for structure within hotspot!



3.7 M\$US approved by Japan for 500 SD
funding fro FD is seeked fo in the US
minor European contribution

GCOS = Global COSmic ray observatory



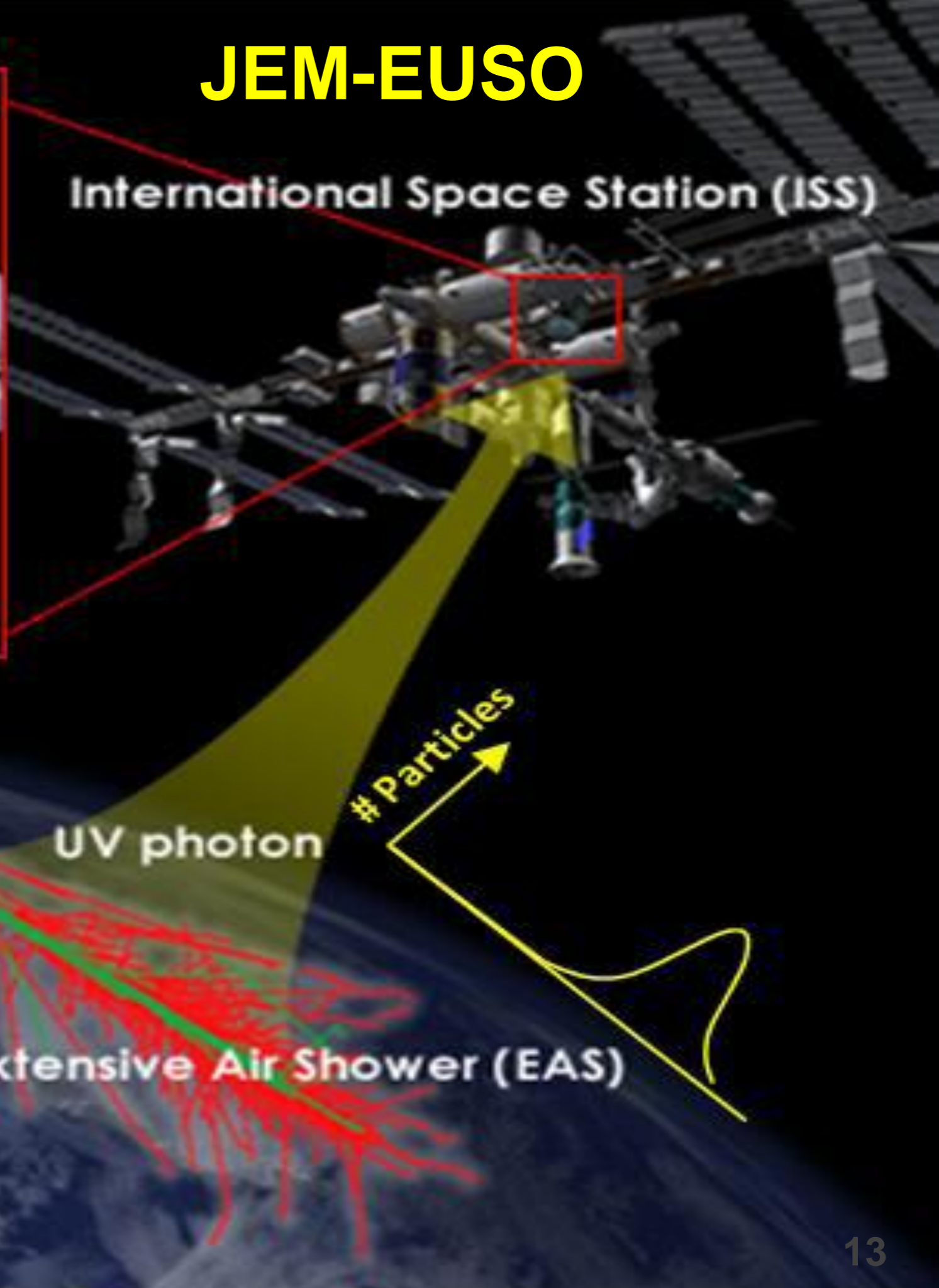
**Helmholtz (D)
large
infrastructure
Roadmap**

p-astronomy with sources

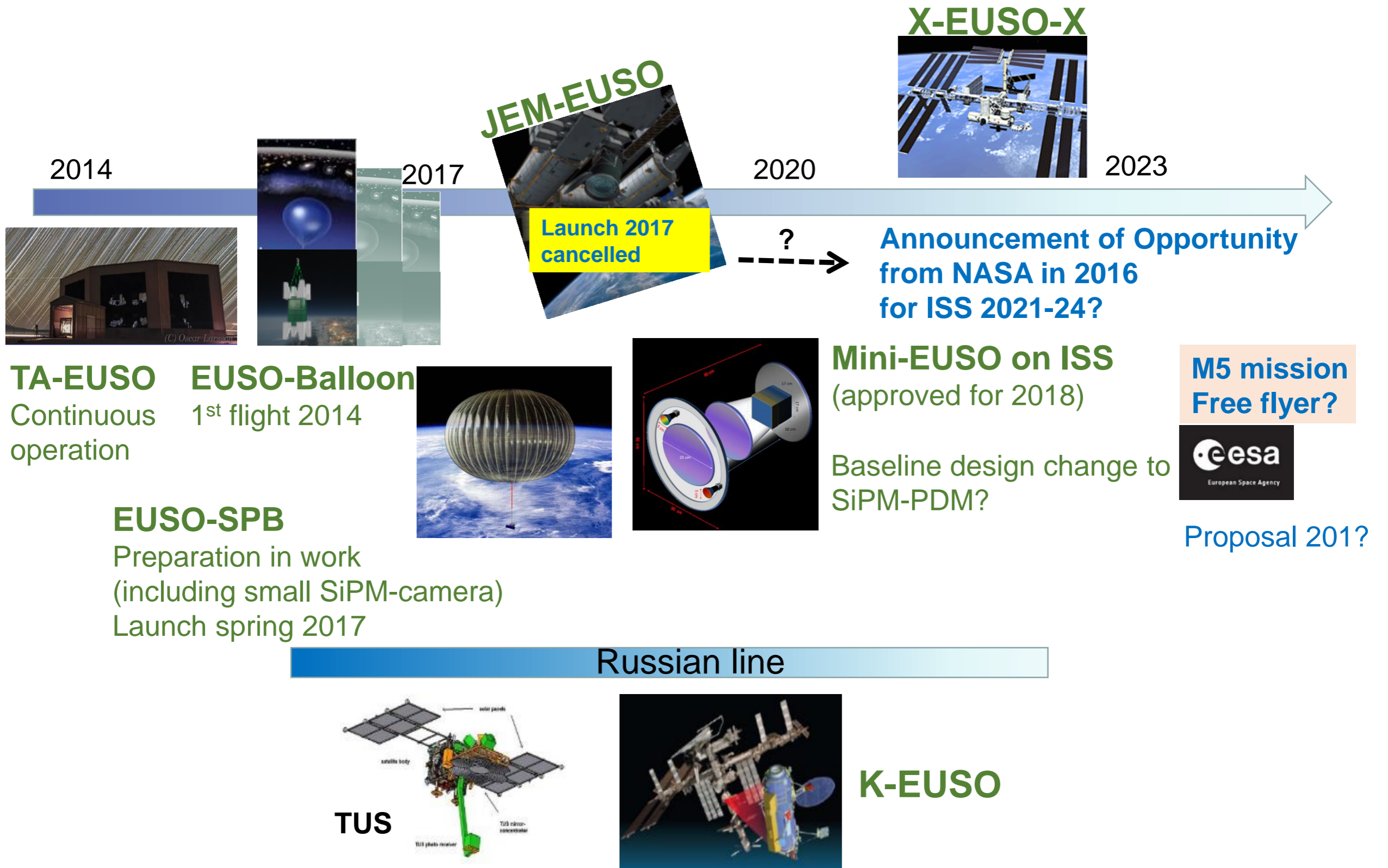
- **Global, few sites, N+S**
- **ca. 90,000 km² (x30 Auger)**
- **Optimal detector for composition-sensitivity**
- **Design in 2020-25**
- **Operation 2025-2050**
- **Cost 390 M€ (120 M€ European contr.)**
- **Operation cost 6 M€/y**

JEM-EUSO

International Space Station (ISS)



Air Shower Observations from Space



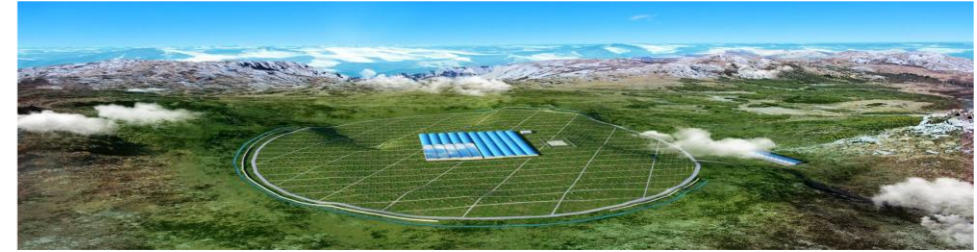
JEM-EUSO collaboration 358 authors; 95 institutions; 16 countries



Further experiments with possible strategic relevance

- **LHAASO**

CR around knee with multi-detector installation
China - with participation of France, Italy



- **TAIGA/Tunka/HiSCORE/Tunka-Taiga-Rex**

CR around knee and up to ankle with multi-detector installation
Russia - with participation of Germany, more?



- **IceCube/IceTop – (Gen2)**

Ice-Cherenkov array on top of IceCube
USA – with important European contribution
Advanced plans for Gen2-surface (veto) array



- **GRAPES**

KASCADE-like operating array at 2300m altitude
India - with participation from Japan



- **KCDC**

KASCADE Cosmic ray Data Centre for public use
Extension to other experiments foreseen (Auger?)



LHAASO



Tunka / Tunka-Rex / HiScore



IceCube / IceTop (-Gen2)



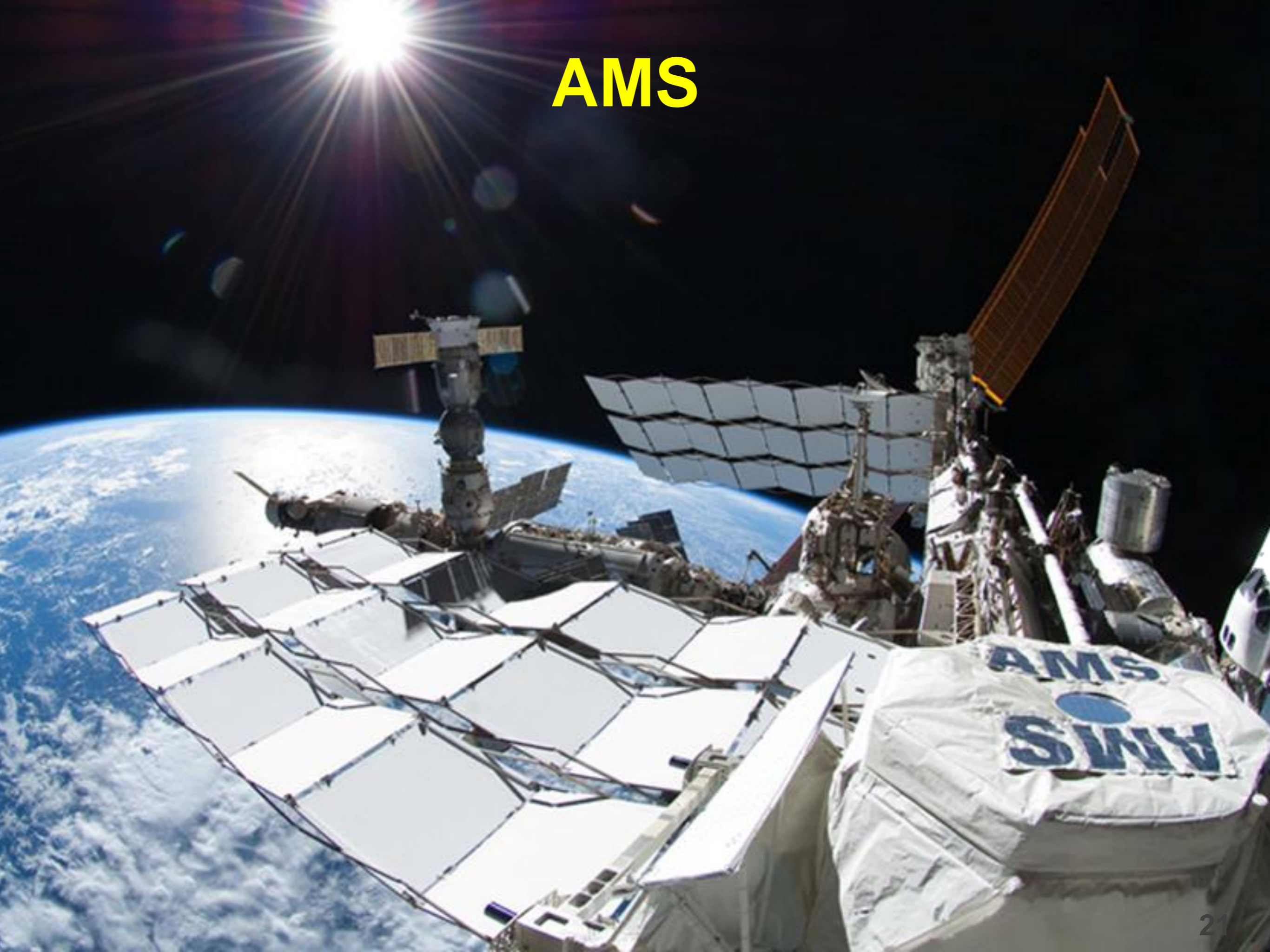
GRAPES



<https://kcdc.ikp.kit.edu>



AMS



SWOT: AugerPrime

Strength

- environmental infrastructure exist
- clear science task
- part of multi-messenger program
- cost efficient
- collaboration (funding agencies) exist
- working groups with TA exist

Weakness

- small statistics at highest energies

Opportunities

- European leadership
- definition of science questions for GCOS
- R&D for next generation experiment

Threats

- uncertainty due to hadronic interaction models

SWOT: all CR

Strength

- global program for all energies
- ready for multi-messenger
- Global use of new large infrastructures
- complementary with LHC
- close community (for air-shower observations)

Weakness

- mostly non-European dominated
- small visibility

Opportunities

- cost effective European contribution
- Participation / driving hardware R&D
- ready / preparing public data dissemination

Threats

- excluded from R&D
- missing visible European participation

APPEC Roadmap for high-energy cosmic-rays

Pierre Auger Observatory is key experiment

- **Full support of Upgrade Strategy (AugerPrime)**
- cost estimation is rational
- R&D for GCOS

JEM-EUSO like detector: situation unclear, test experiments running!

- **Support of further design and R&D studies**
- if, than funding by space agencies (?)

Other experiments with small European contribution

- **Mentioning and support of continuation of R&D**
- **Support usage of Infrastructures**

Low energy (space/balloon) cosmic ray investigation?

- AMS analysis / ISS-CREAM? Funding by space agencies?

Data dissemination (open access to scientific data)

- **Need more room and efforts in next decade**
- common funding?